

## An SU(3) symmetry for light neutrinos

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**Abstract.** It is proposed that light neutrinos form a triplet in a global SU(3) symmetry in the mass eigenstate basis. Assuming that the SU(3) symmetry is broken in the direction  $(-a\lambda_3 + \frac{b}{\sqrt{3}}\lambda_8)$ , and after going to the flavor basis, we predict the atmospheric mixing angles  $\sin^2 \theta_{23} = 0.5$  and  $\sin \theta_{13} = 0$ , if  $\nu_\mu - \nu_\tau$  symmetry is assumed. In the flavor basis, the diagonal part of the matrix coefficient of  $b$  (the dominant part) is found to transform like  $(\lambda_3 + \frac{1}{\sqrt{3}}\lambda_8)$ . Imposing the same condition on the matrix coefficient of  $a$  fixes the solar mixing angle,  $\sin^2 \theta_{12} = \frac{1}{3}$ . The implications for neutrinoless double beta decay are discussed.

There is compelling evidence (for a recent discussion and many references, see [1]) that neutrinos change flavor, have non-zero masses and that the neutrino mass eigenstates are different from the weak eigenstates. As such they undergo oscillations.

All neutrino data [1] with the exception of the LSND anomaly [2] are explained by three neutrino flavor oscillations with mass squared differences and mixing angles having the following values [3]:

$$\begin{aligned} \Delta m_{\text{solar}}^2 &= \Delta m_{12}^2 = (8.1 \pm 1.0) \times 10^{-5} \text{ eV}^2, \\ \sin^2 \theta_{12} &= 0.30 \pm 0.08, \\ \Delta m_{\text{atm}}^2 &= |\Delta m_{13}^2| \simeq |\Delta m_{23}^2| \\ &= (2.2 \pm 1.1) \times 10^{-3} \text{ eV}^2, \\ \sin^2 \theta_{23} &= 0.50 \pm 0.18, \\ \sin^2 \theta_{13} &\leq 0.047. \end{aligned} \quad (1)$$

The above mixing pattern is in conformity with the “bi-tri-maximal” scheme, first discussed in [4, 5]. The neutrino mixing angles are defined by the lepton mixing matrix [6]

$$\begin{pmatrix} \nu_e \\ \nu_\mu \\ \nu_\tau \end{pmatrix} = U \begin{pmatrix} \nu_1 \\ \nu_2 \\ \nu_3 \end{pmatrix}. \quad (2)$$

The matrix  $U$  is conveniently parametrized by the three mixing angles  $\theta_{12}$ ,  $\theta_{13}$  and  $\theta_{23}$  and the three complex phases, two of which are the so-called Majorana phases. We put all phases to zero.

The purpose of this paper is to study the implications of SU(3) symmetry for light neutrinos in the mass eigen-

state basis. SU(3) family symmetry has previously been used [7, 8] in a more fundamental way than attempted here. Our aim is modest. We show that if the SU(3) symmetry for the neutrino mass eigenstates is broken in the direction  $(-a\lambda_3 + \frac{b}{\sqrt{3}}\lambda_8)$ , and if we go to the flavor basis by the unitary transformation given in (2), the atmospheric mixing angles are predicted to be  $\sin^2 \theta_{23} = 0.5$  and  $\sin \theta_{13} = 0$ , if the  $\nu_\mu - \nu_\tau$  symmetry is assumed. For recent use of  $\nu_e \leftrightarrow \nu_\mu$  symmetry in other contexts, see for example, [9–11]. This symmetry is inspired by the experimental observation of a near maximal atmospheric mixing angle and a small upper limit on  $\theta_{13}$ , implying the interesting possibility that there may be an approximate  $\mu \leftrightarrow \tau$  symmetry in the neutrino sector [12–22]. Its deeper origin is not yet known. Such a symmetry has also interesting implications in leptogenesis; see, for example, [23–25].

It is found that in a flavor basis, the diagonal part of the matrix coefficient of  $b$  (dominant part) exhibits an interesting property, namely that it transforms like  $(\lambda_3 + \frac{1}{\sqrt{3}}\lambda_8)$ . If we impose the same condition on the matrix coefficient of  $a$  (non-leading part) we also predict the solar mixing angle  $\sin^2 \theta_{12} = \frac{1}{3}$ . In our approach the absolute mass of the neutrino in the SU(3) limit,  $m_0$ , is not constrained by neutrino oscillation data. If the WMAP constraint on the neutrino mass,  $\sum m_i < (0.4\text{--}0.7) \text{ eV}$ , is used, the effective double beta decay mass  $m_{ee} < (0.13\text{--}0.23) \text{ eV}$ .

We assume that the mass eigenstates  $(\nu_1, \nu_2, \nu_3)$  form an SU(3) triplet, so that, in the SU(3) limit, they have the common mass  $m_0$ . Since the neutrino mass matrix in the mass eigenstate basis has to be diagonal and the only diagonal matrices available are  $\lambda_3$  and  $\lambda_8$ , the most general form of symmetry breaking is provided by  $(-a\lambda_3 + \frac{b}{\sqrt{3}}\lambda_8)$ ,

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so that in the basis  $(\nu_1, \nu_2, \nu_3)$  we have

$$\mathcal{M} = m_0 I + \left( -a\lambda_3 + \frac{b}{\sqrt{3}}\lambda_8 \right), \quad (3)$$

$$\begin{aligned} m_1 &= m_0 - a + \frac{b}{3} = m - a, \\ m_2 &= m_0 + a + \frac{b}{3} = m + a, \\ m_3 &= m_0 - \frac{2b}{3} = m - b, \end{aligned} \quad (4)$$

where  $m = m_0 + \frac{b}{3}$  and  $|a|, |b| \ll m_0$ .

Thus, we have

$$\begin{aligned} \Delta m_{12}^2 &\simeq 4ma, \\ |\Delta m_{23}^2| &\simeq 2m(a+b) + (b^2 - a^2), \\ |\Delta m_{13}^2| &\simeq 2m(b-a) + (b^2 + a^2). \end{aligned} \quad (5)$$

The data require that  $a \ll b$ , so that  $m_3 \ll m_1 \leq m_2$  (inverted hierarchy) and

$$|\Delta m_{23}^2| \simeq |\Delta m_{13}^2| \simeq 2mb. \quad (6)$$

We now go to the flavor basis  $(\nu_e, \nu_\mu, \nu_\tau)$  by using (2), [ $s_1 = \sin \theta_{23}$ ,  $c_1 = \cos \theta_{23}$  and  $s_2 \equiv \sin \theta_{13}$ ]. The neutrino mass matrix in the flavor basis then is

$$M_\nu = m_0 I + M,$$

with

$$\begin{aligned} M_{11} &= \frac{b}{3} (1 - 3s_2^2) - ac_2^2 \cos 2\theta_{12}, \\ M_{22} &= \frac{b}{3} (1 - 3s_1^2 c_2^2) \\ &\quad - a [\cos 2\theta_{12} (-c_1^2 + s_1^2 s_2^2) + \sin 2\theta_1 \sin 2\theta_{12} s_2], \\ M_{33} &= \frac{b}{3} (1 - 3c_1^2 c_2^2) \\ &\quad - a [\cos 2\theta_{12} (-s_1^2 + c_1^2 s_2^2) - \sin 2\theta_1 \sin 2\theta_{12} s_2], \\ M_{12} &= -bc_2 s_2 s_1 + ac_2 [c_1 \sin 2\theta_{12} + s_1 s_2 \cos 2\theta_{12}], \\ M_{13} &= -bc_2 s_2 c_1 - ac_2 [s_1 \sin 2\theta_{12} - c_1 s_2 \cos 2\theta_{12}], \\ M_{23} &= -bc_1 s_1 c_2^2 \\ &\quad - a [s_1 c_1 \cos 2\theta_{12} (1 + s_2^2) - \sin 2\theta_{12} s_2 (s_1^2 - c_1^2)]. \end{aligned} \quad (7)$$

Imposing  $\nu_\mu \leftrightarrow \nu_\tau$  symmetry, we get

$$s_2 = 0, \quad c_1 = -s_1 = \frac{1}{\sqrt{2}}.$$

Thus,  $M$  reduces to

$$\begin{aligned} &\frac{b}{2} \begin{pmatrix} \frac{2}{3} & 0 & 0 \\ 0 & -\frac{1}{3} & 1 \\ 0 & 1 & -\frac{1}{3} \end{pmatrix} \\ &- \frac{a}{2} \begin{pmatrix} 2 \cos 2\theta_{12} & -\sqrt{2} \sin 2\theta_{12} & -\sqrt{2} \sin 2\theta_{12} \\ -\sqrt{2} \sin 2\theta_{12} & -\cos 2\theta_{12} & -\cos 2\theta_{12} \\ -\sqrt{2} \sin 2\theta_{12} & -\cos 2\theta_{12} & -\cos 2\theta_{12} \end{pmatrix}. \end{aligned} \quad (8)$$

It is interesting to note that the diagonal part of the matrix coefficient of  $\frac{b}{2}$  transforms as  $\lambda_3 + \frac{1}{\sqrt{3}}\lambda_8$ , like the electric charge operator of the  $u$ ,  $d$  and  $s$  quarks. If we require the same for the matrix coefficient of  $-a/2$ , we obtain

$$\cos 2\theta_{12} = \frac{1}{3},$$

giving

$$\sin^2 \theta_{12} = \frac{1}{3}.$$

This is consistent with its experimental value given in (1). Thus, the neutrino mass matrix in the flavor basis is

$$\begin{aligned} M_\nu &= m_0 I + \frac{b}{2} \begin{pmatrix} 2/3 & 0 & 0 \\ 0 & -1/3 & 1 \\ 0 & 1 & -1/3 \end{pmatrix} \\ &- \frac{a}{2} \begin{pmatrix} 2/3 & -4/3 & -4/3 \\ -4/3 & -1/3 & -1/3 \\ -4/3 & -1/3 & -1/3 \end{pmatrix}. \end{aligned} \quad (9)$$

The data give

$$\begin{aligned} \frac{a}{b} &\simeq \frac{1}{2} \frac{\Delta m_{12}^2}{|\Delta m_{23}^2|} \simeq 1.8 \times 10^{-2} \\ \frac{b}{m_0} &\simeq \frac{1}{2} \frac{|\Delta m_{23}^2|}{m_0^2} \simeq (1.1) \times 10^{-3} \frac{\text{eV}^2}{m_0^2}. \end{aligned} \quad (10)$$

$m_0$  is not constrained by the oscillation data. However,  $m_0$  is constrained by the WMAP data,  $\sum m_i < (0.4-0.7) \text{ eV}$ .

Thus, taking

$$m_0 \simeq 0.1 \text{ eV},$$

we have

$$\frac{b}{m_0} \approx 10^{-1}. \quad (11)$$

The SU(3) symmetry thus makes sense, as the symmetry breaking parameter is small.

Finally, for neutrinoless double  $\beta$  decay, the effective double beta decay mass  $\langle m_{ee} \rangle$  is predicted to be [ $\sigma_1$  and  $\sigma_2$  are Majorana phases,  $\theta_{13} = 0$ ] [26]

$$\begin{aligned} m_{ee} &= \left| |m_1| \cos^2 \theta_{12} e^{-2i\sigma_1} + |m_2| \sin^2 \theta_{12} e^{-2i\sigma_2} \right| \\ &\simeq \frac{m_0}{3} < m_{ee} < m_0. \end{aligned} \quad (12)$$

Using the WMAP limit on  $m_0$ , we get

$$m_{ee} \leq (0.13-0.23) \text{ eV}.$$

In summary, a global SU(3) symmetry for the neutrino mass eigenstates and its breaking along the direction  $(-a\lambda_3 + \frac{b}{\sqrt{3}}\lambda_8)$  with  $a \ll b$ , together with the  $(\nu_\mu \leftrightarrow \nu_\tau)$  symmetry in the flavor basis and the requirement that the diagonal part of the neutrino mass matrix  $(M_\nu - m_0 I)$  transforms as  $(\lambda_3 + \frac{1}{\sqrt{3}}\lambda_8)$ , can explain the data in (1).

Furthermore, together with the WMAP constraint on the neutrino mass, an effective double beta decay mass  $m_{ee}$  is predicted.

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